

FUSION

the energy of the 3rd millenium



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PART I: PEOPLE AND ENERGY



What is it, the energy?

- **in physics:** a scalar quantity subject to conservation law, measured originally as a potential to do work (in Joules) leading to many consequences in theory, including

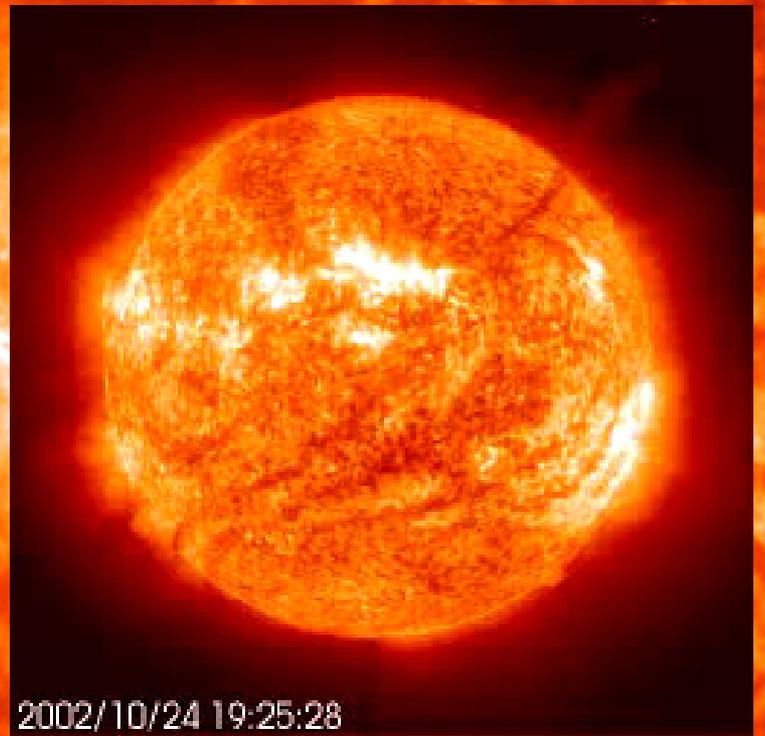
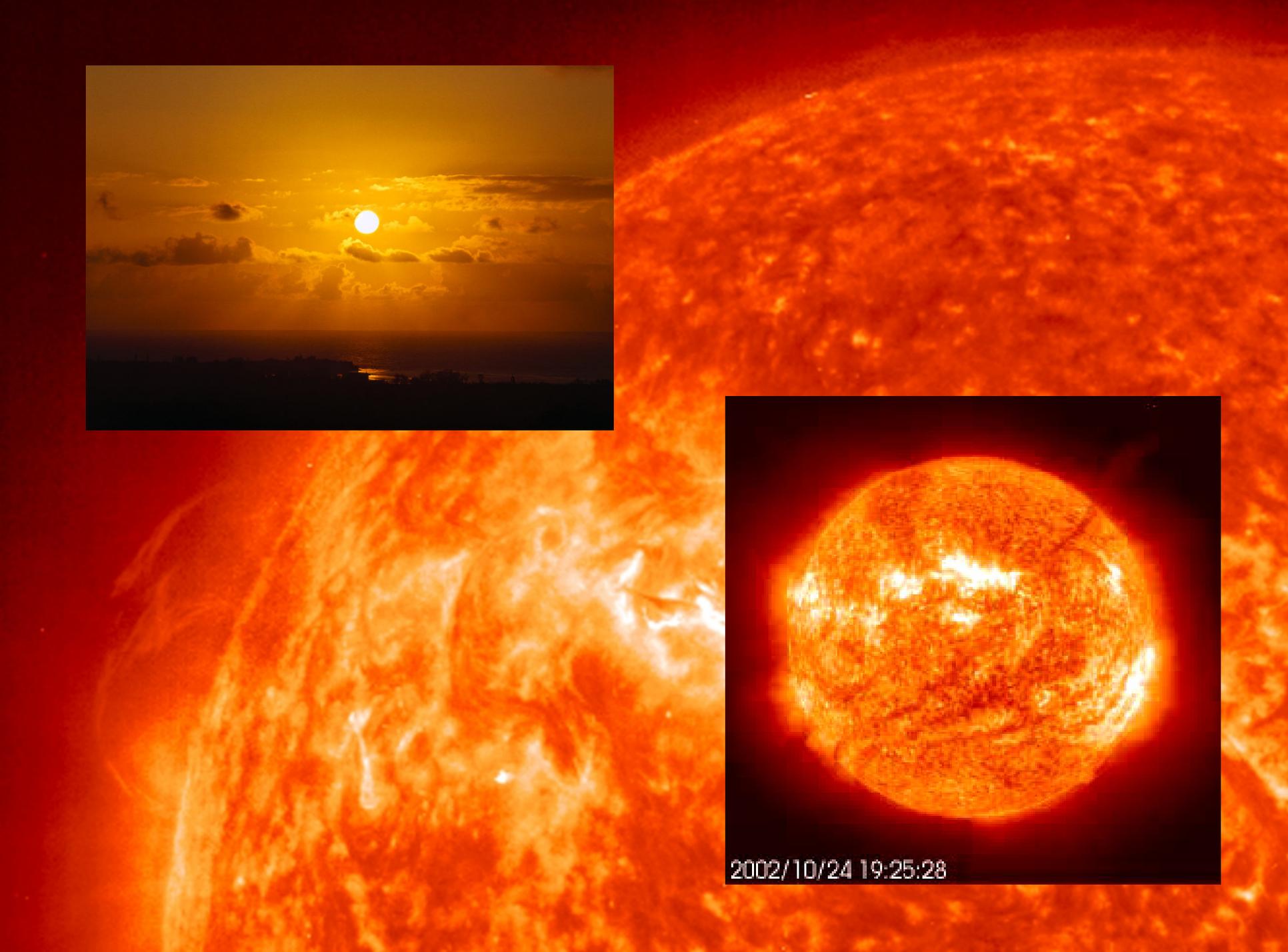
$$E = mc^2$$

- **in common sense:**

1. vitality i.e. capability to act, or even a spirituality in esotericism, life force (beyond the scope of this talk)
2. “the resource that is spent when something is done”, often measured (in Joules, kWh or calories) i.e. its **form**, “**quality**” is much more important than the conservation law.

*energy has **high quality** if its form is easily convertible into other forms (compare electricity – heat – rest energy)*

Where does it come from, the energy?



Solar deity



Egypt:: God Ra



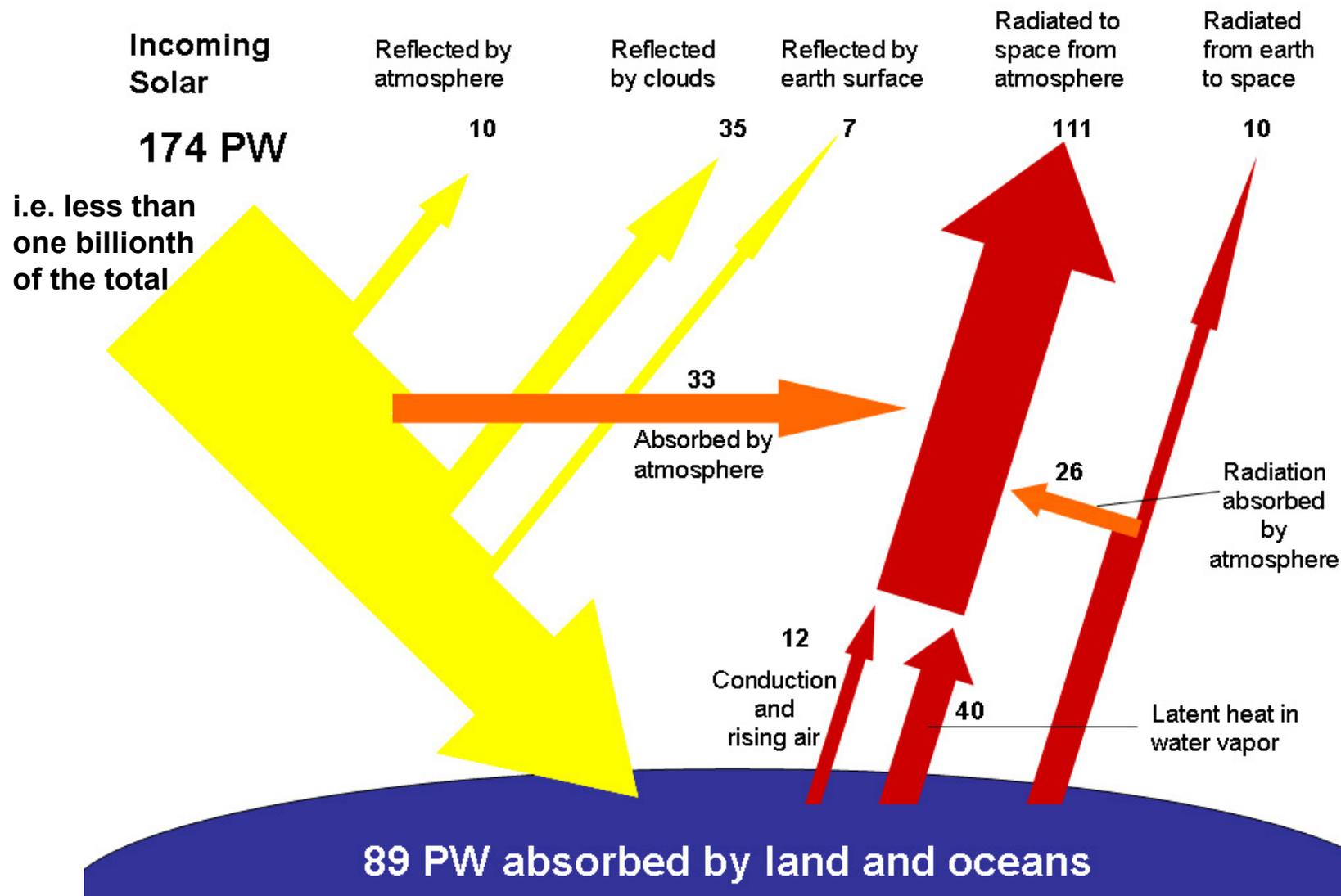
Slavic paganism: God Radegast



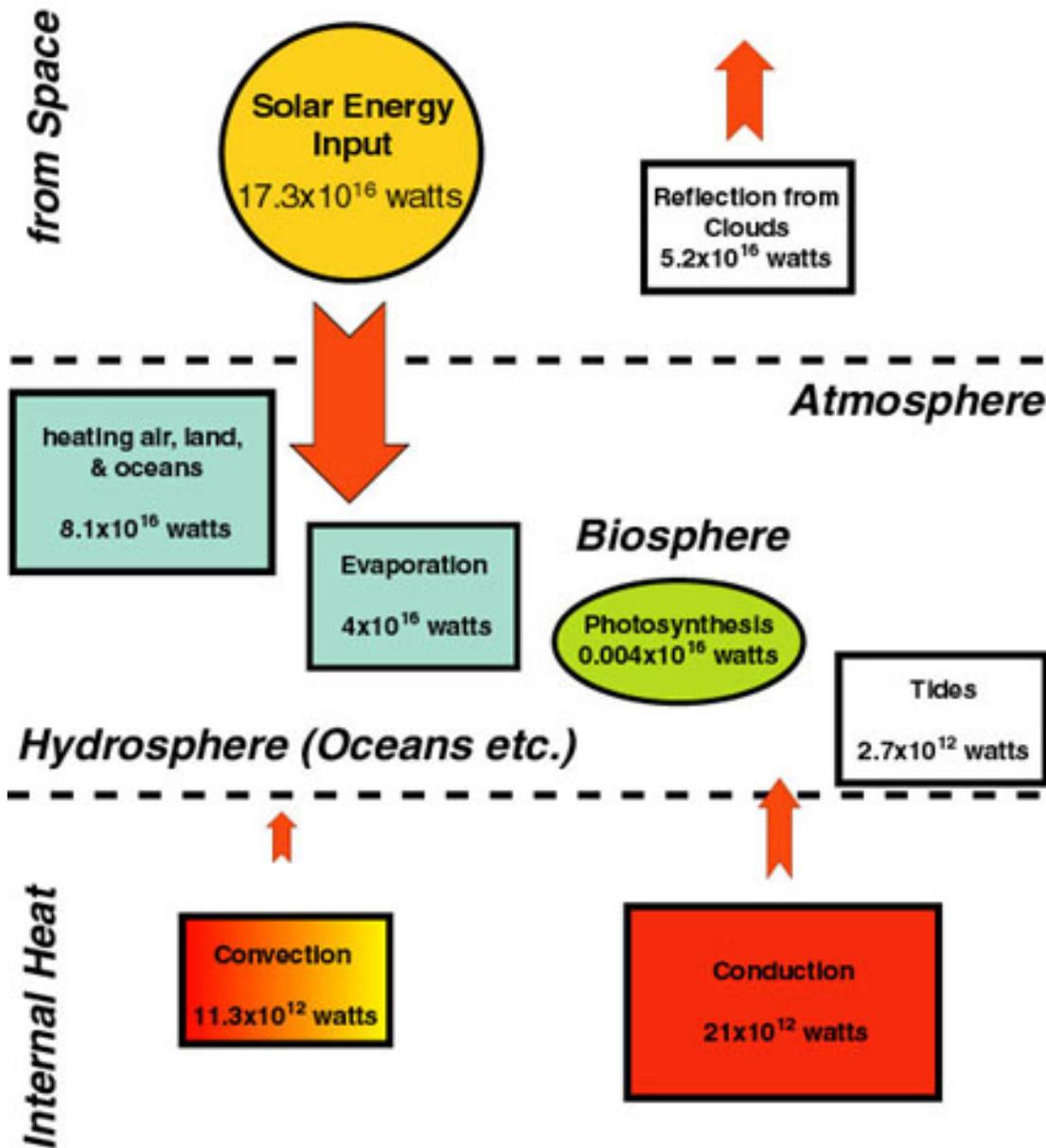
Aztec: God Tonatiuh

Solar power at the Earth

Total power of the Sun is $4 \cdot 10^{26} \text{ W} = 4 \cdot 10^{11} \text{ PW}$



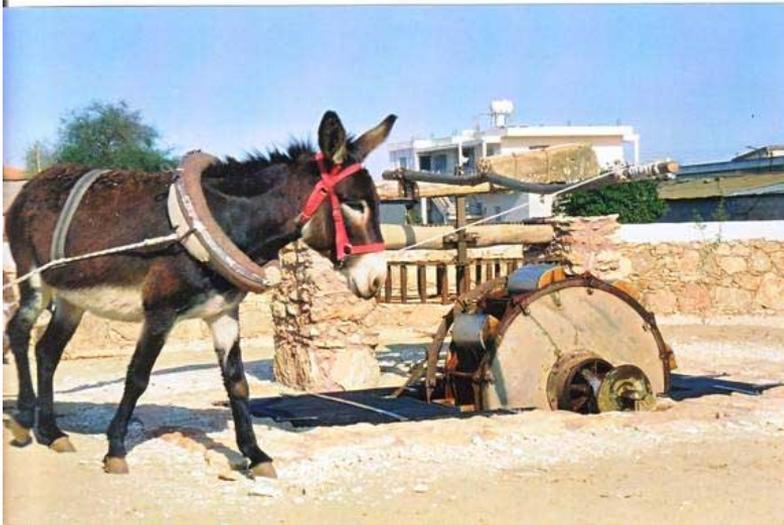
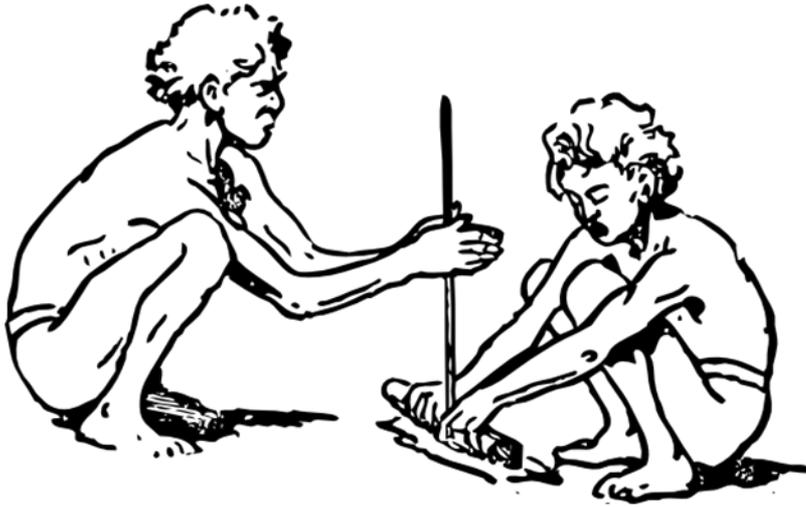
The Earth's Energy Budget



Photosynthesis
 4.10^{13} W = 40 TW

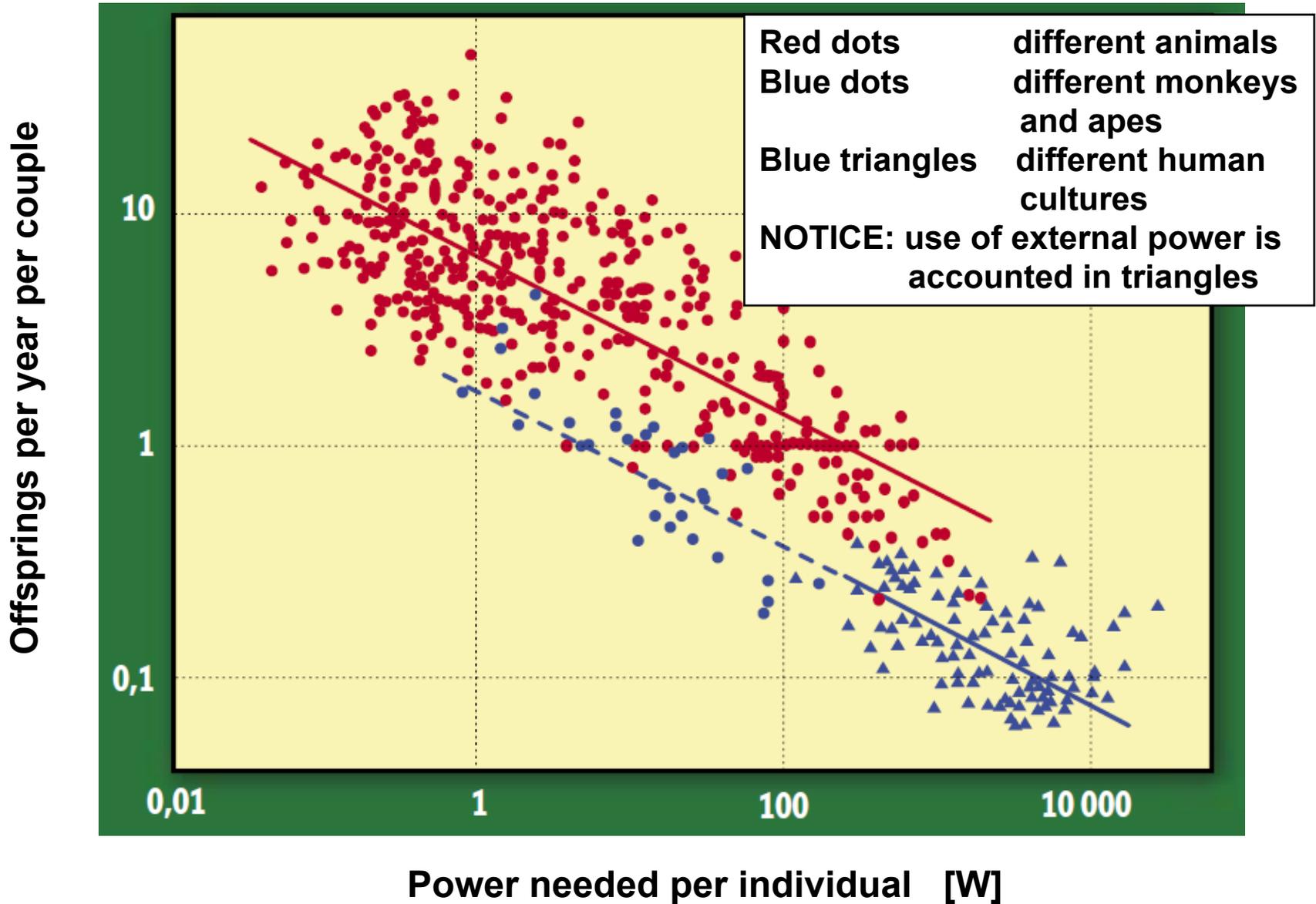
Humankind
total power 16 TW
in renewables ~ 1 TW
in electricity ~ 2 TW

Why do we need energy?



Will we always ask for more energy? No, now rich people ask for better quality.

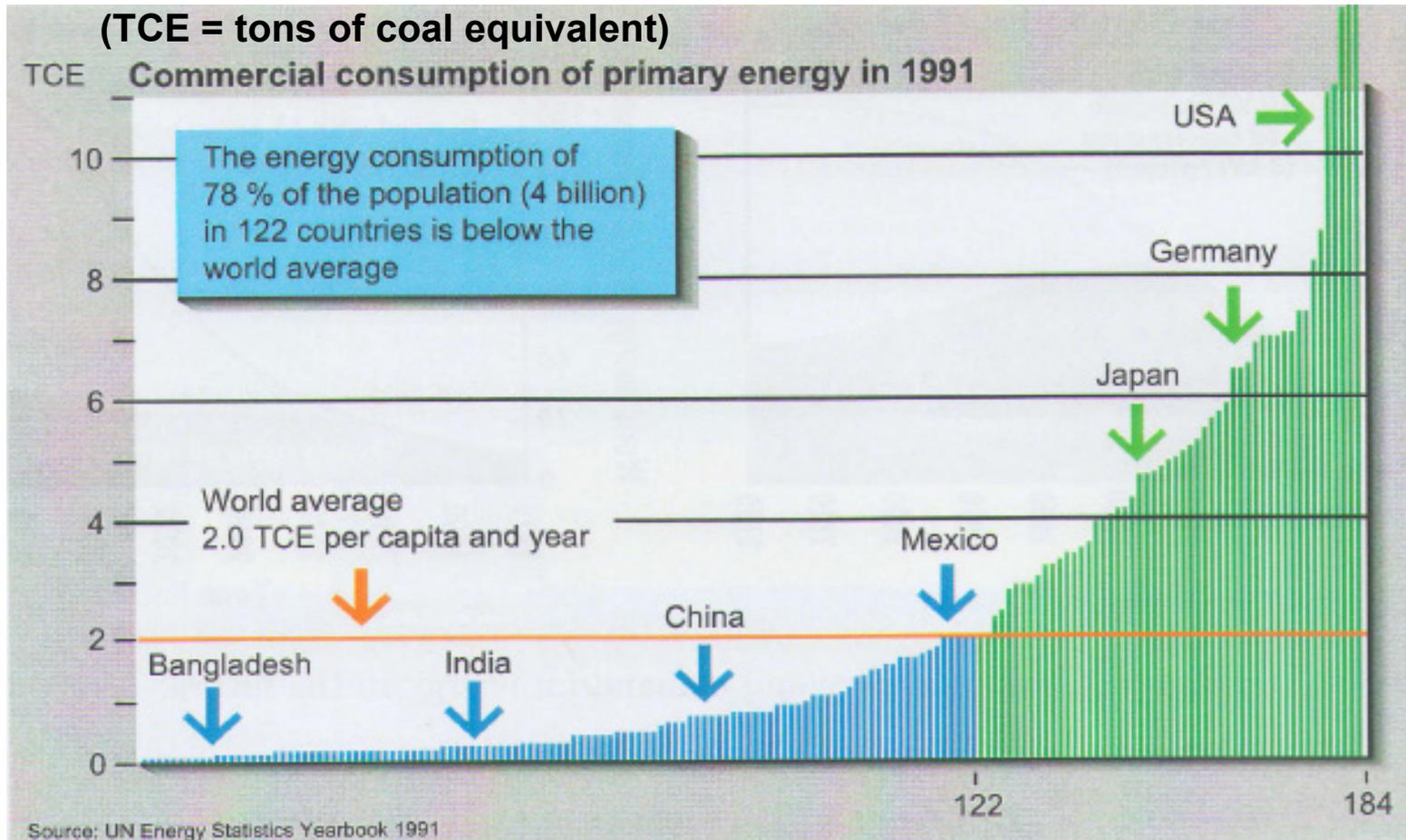
Time for break: a weird link of human and animals



The energy needs of the world

Global energy consumption will increase due to

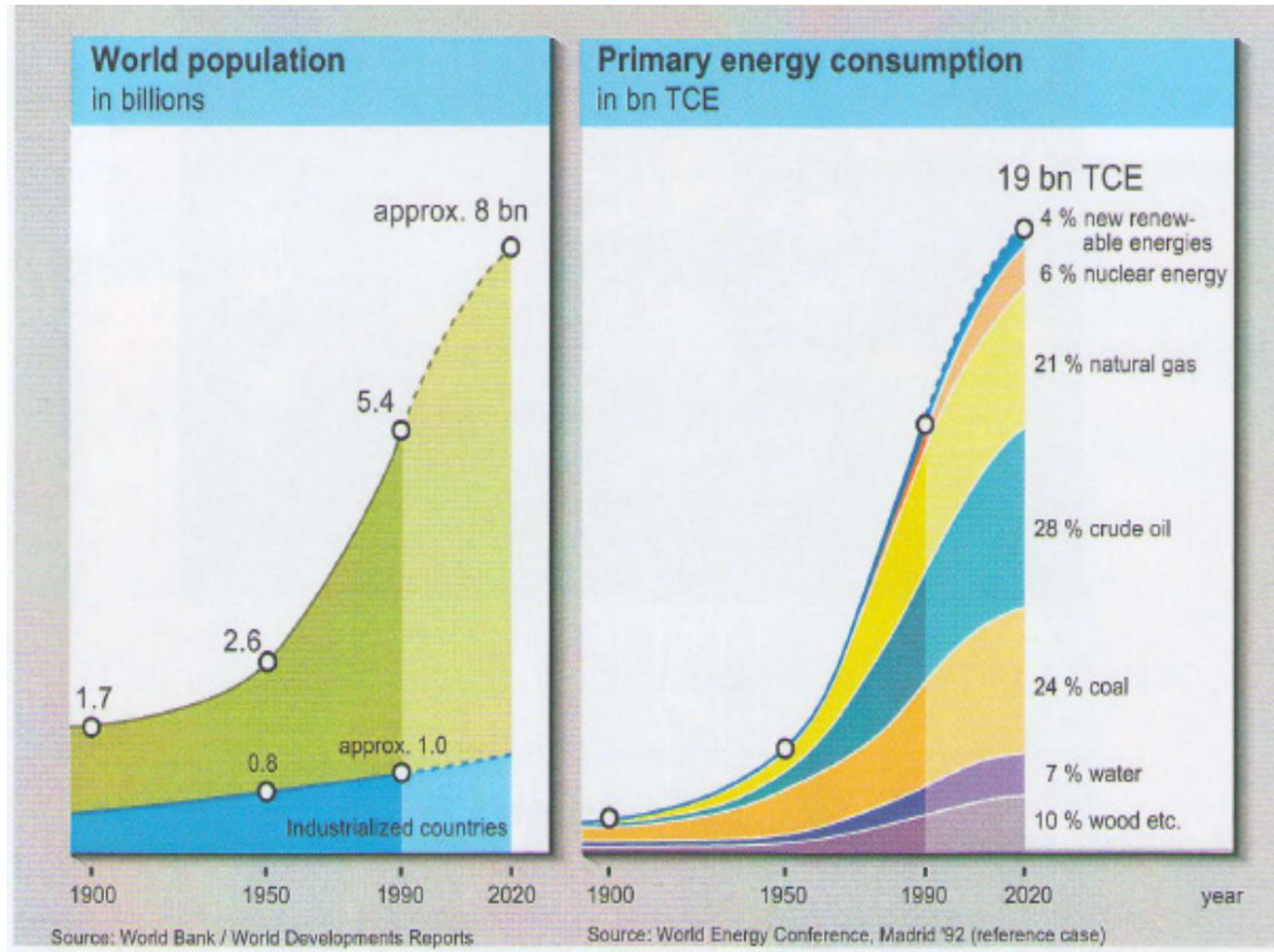
- population growth
- economic progress in (some) poor countries



acceptable quality of life ~ 3 TCE on average.

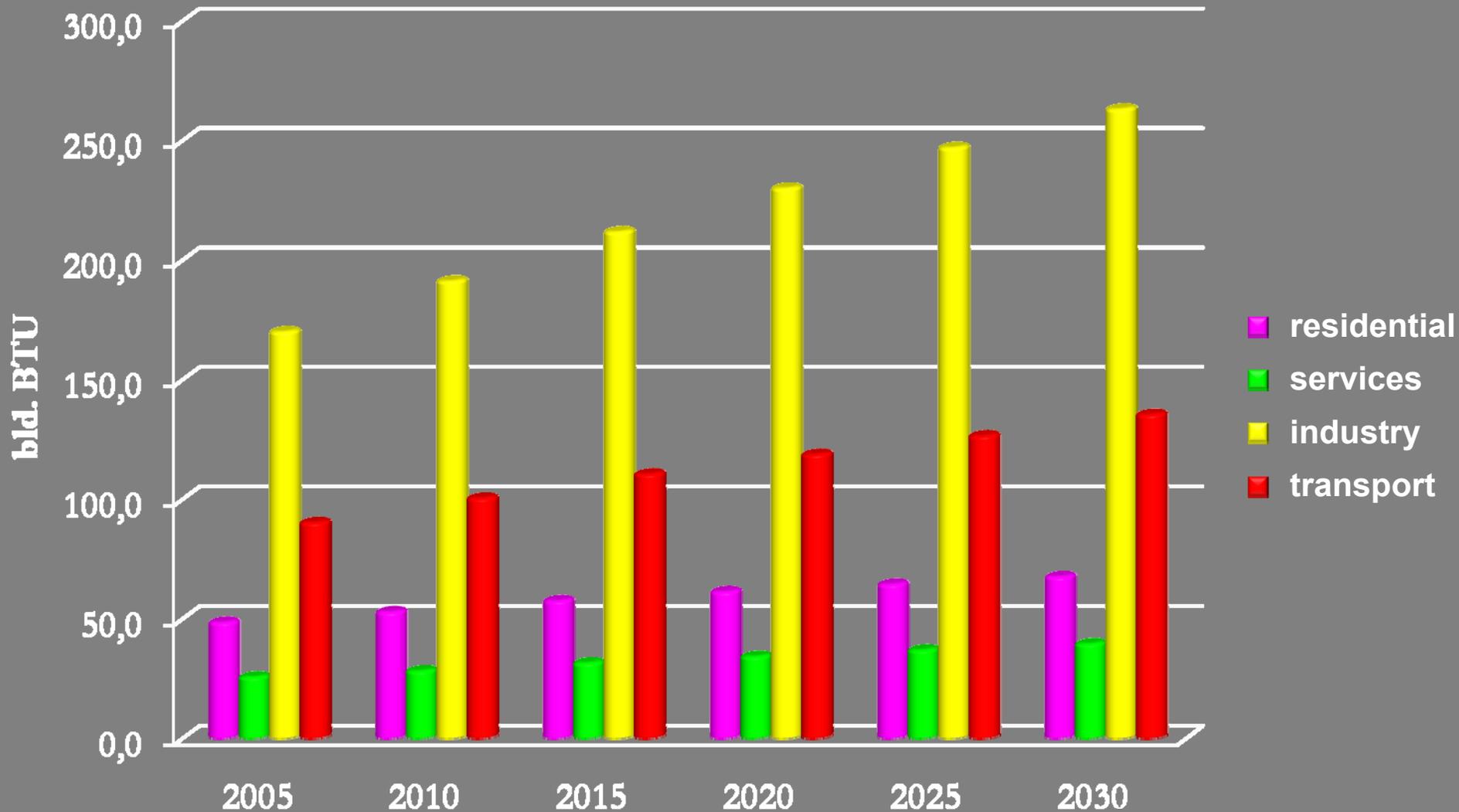
People need more energy and cleaner energy

IEA predicts the global energy consumption will double by 2045.



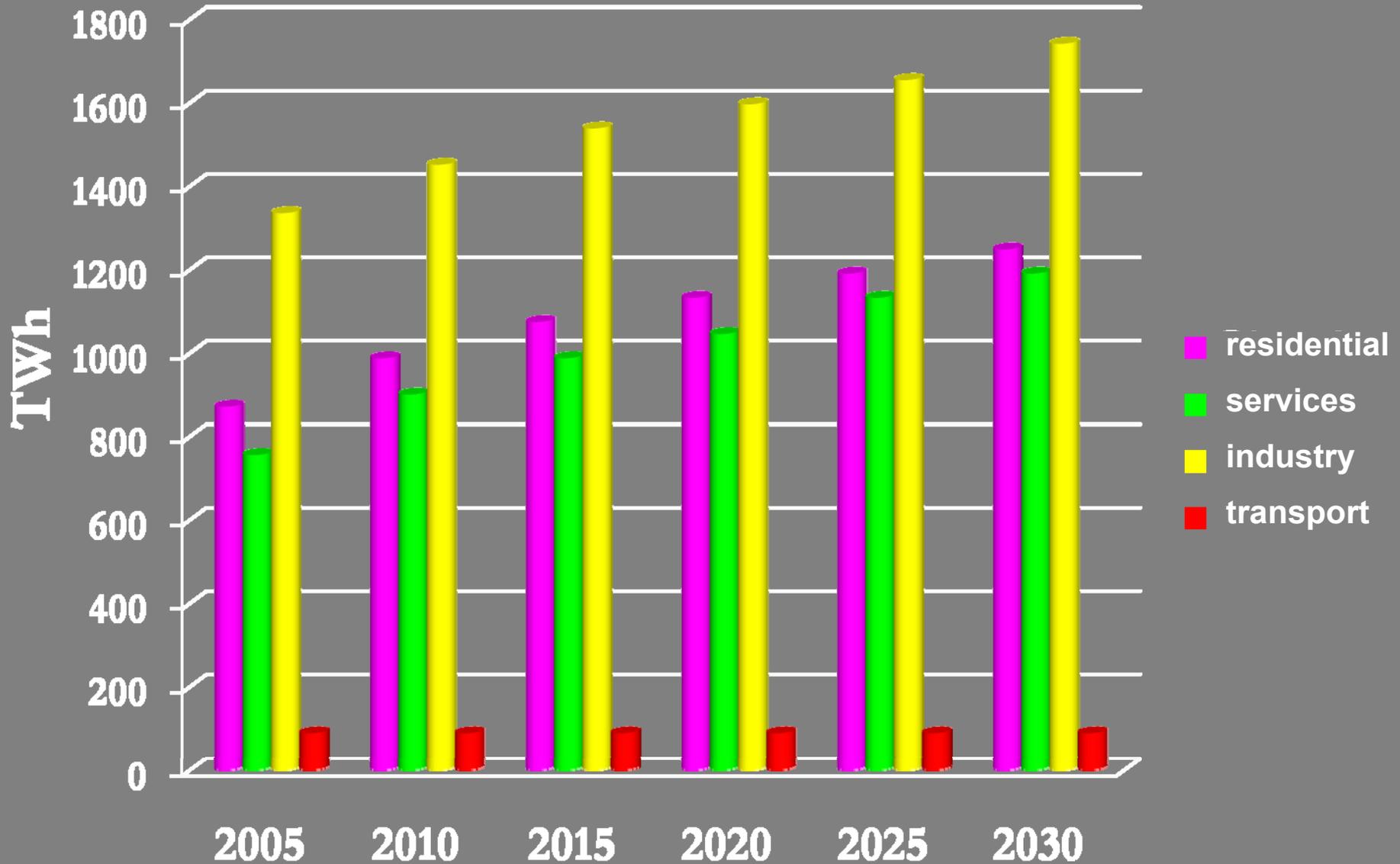
Don't give up! Remind that we still consume much less than the rest of living nature. The challenge is to get a sustainable power source.

Consumption of primary energy



BTU = British Thermal Unit ~ 1055 J

Consumption of electricity



Fossil fuels: 3/4 of our energy consumptions

Fossil fuels took millions of years to form. Future generations will not appreciate that we depleted their resources.

And contributed to...

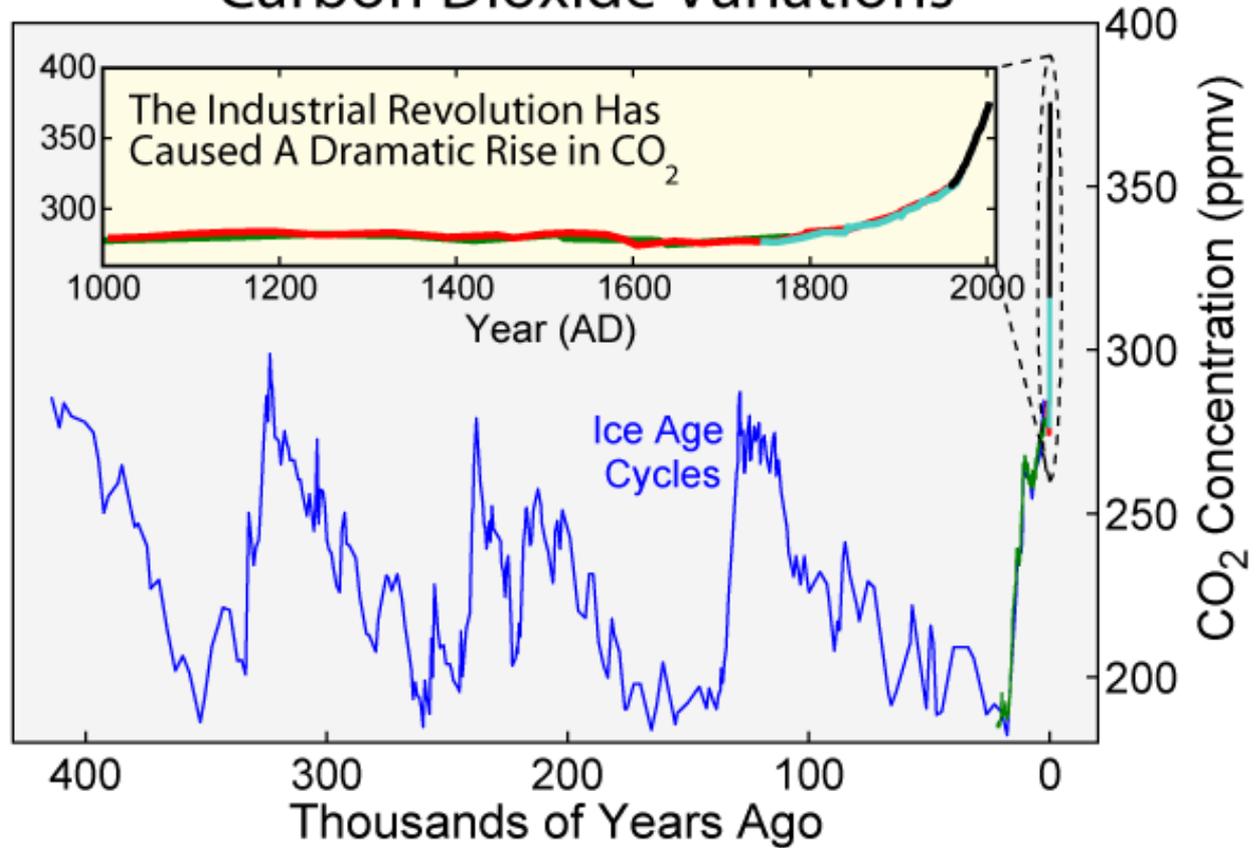


- **climate change**
- **geopolitical and cultural conflicts**
- **local pollution**

Fossil fuels: Changing the air



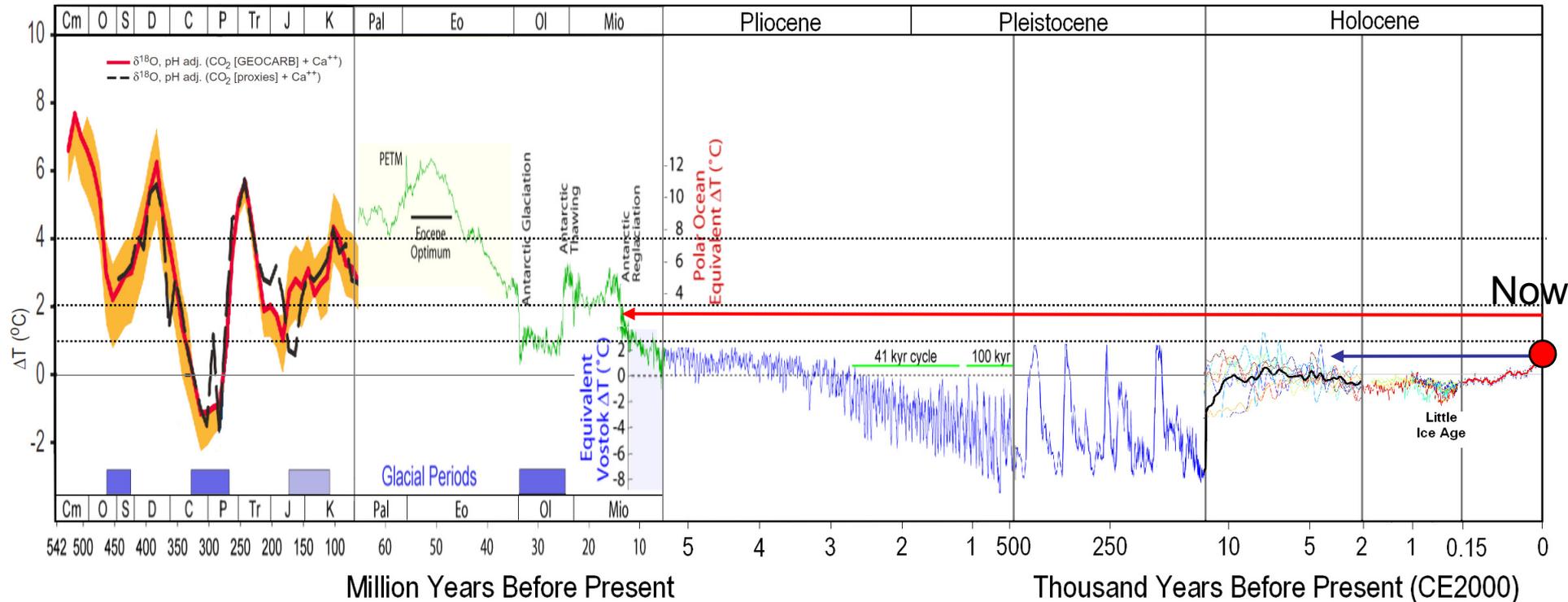
Carbon Dioxide Variations



Is CO₂ safe??

Greenhouse effect and climate change

Temperature of Planet Earth



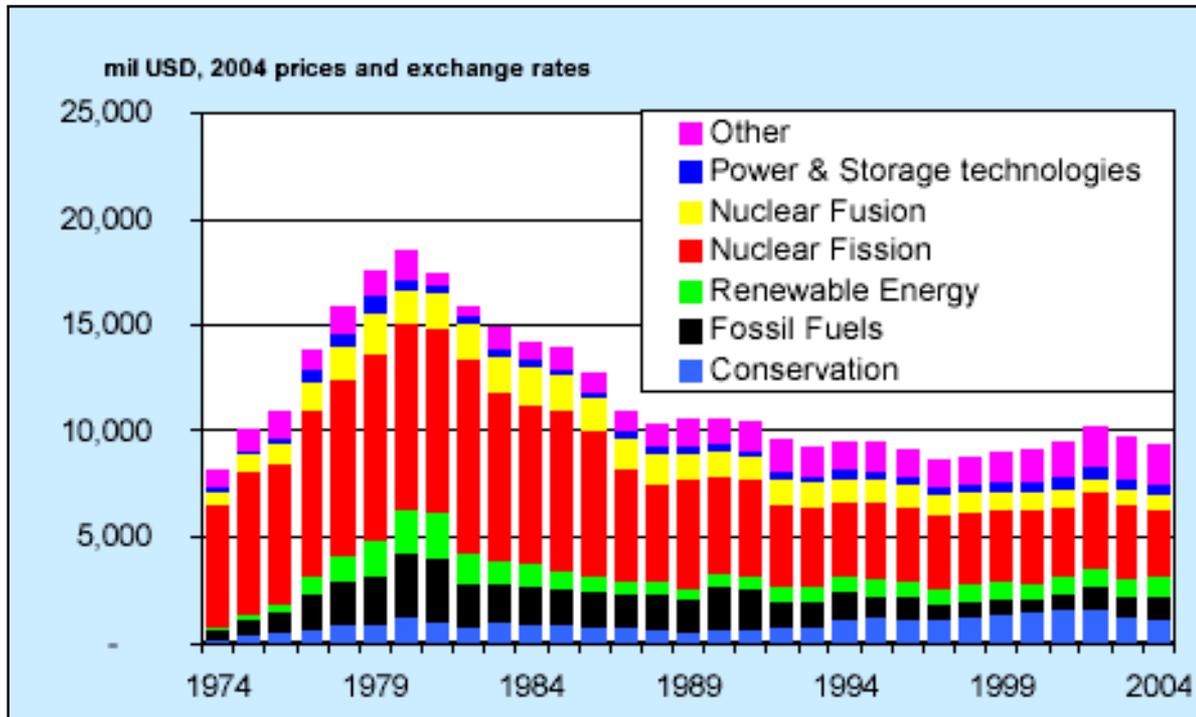
- Global surface temperature increased **$0.74 \pm 0.18^\circ\text{C}$** during the last century.
- Projections indicate that the global surface temperature will probably rise a further 1.1 to 6.4°C during the 21st century,
- Increasing global temperature will cause sea levels to rise.

World energy outlook

„Stern review“

(Independent study for the British government) 2006

If the climate change is due to the human activities, than we should either invest 1% of the GNP into making these activities sustainable, or get prepared for 20% decrease of the GNP.

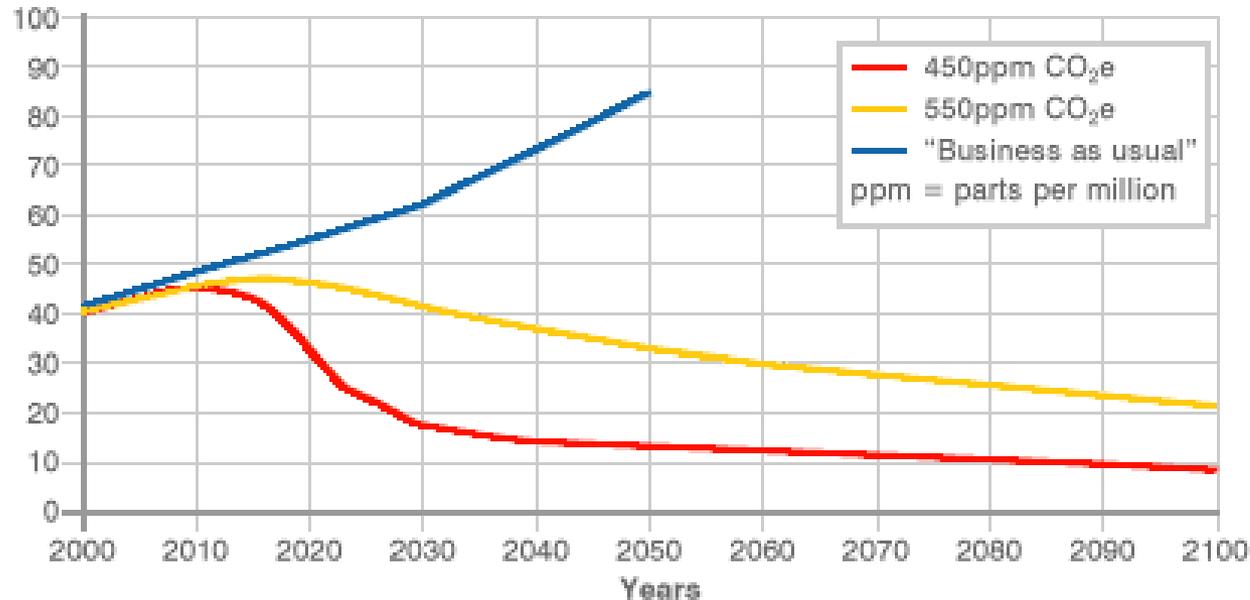


Unfortunately, the leading economies do not invest much into energy research (neither public nor private bodies) due to long term return of the money invested.

Public energy R&D in IEA countries

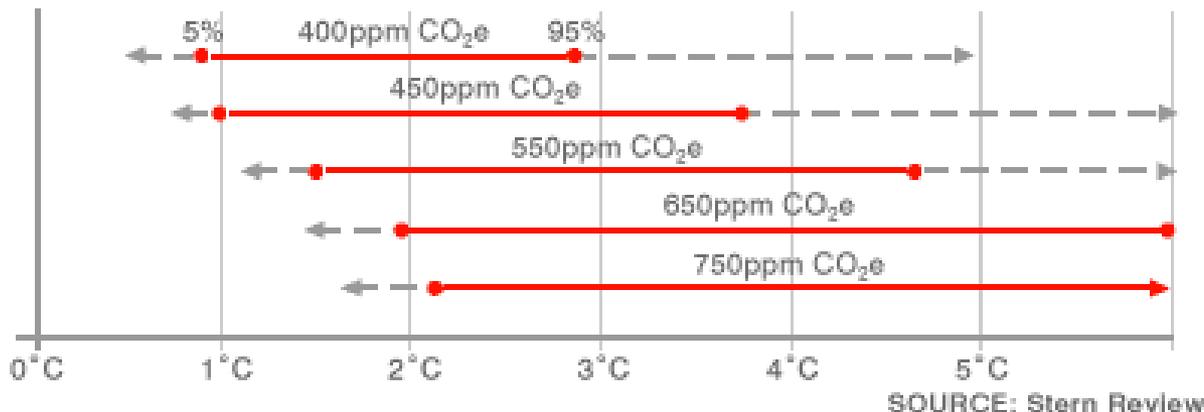
Emission paths to stabilisation

Global Emissions (Gigatonnes of CO₂ equivalent gases per year)



Kyoto protocol 1997
in force since 2005
Mission: based on 1990 data cut 7% emissions of greenhouse gasses incl. CO₂ by 2012

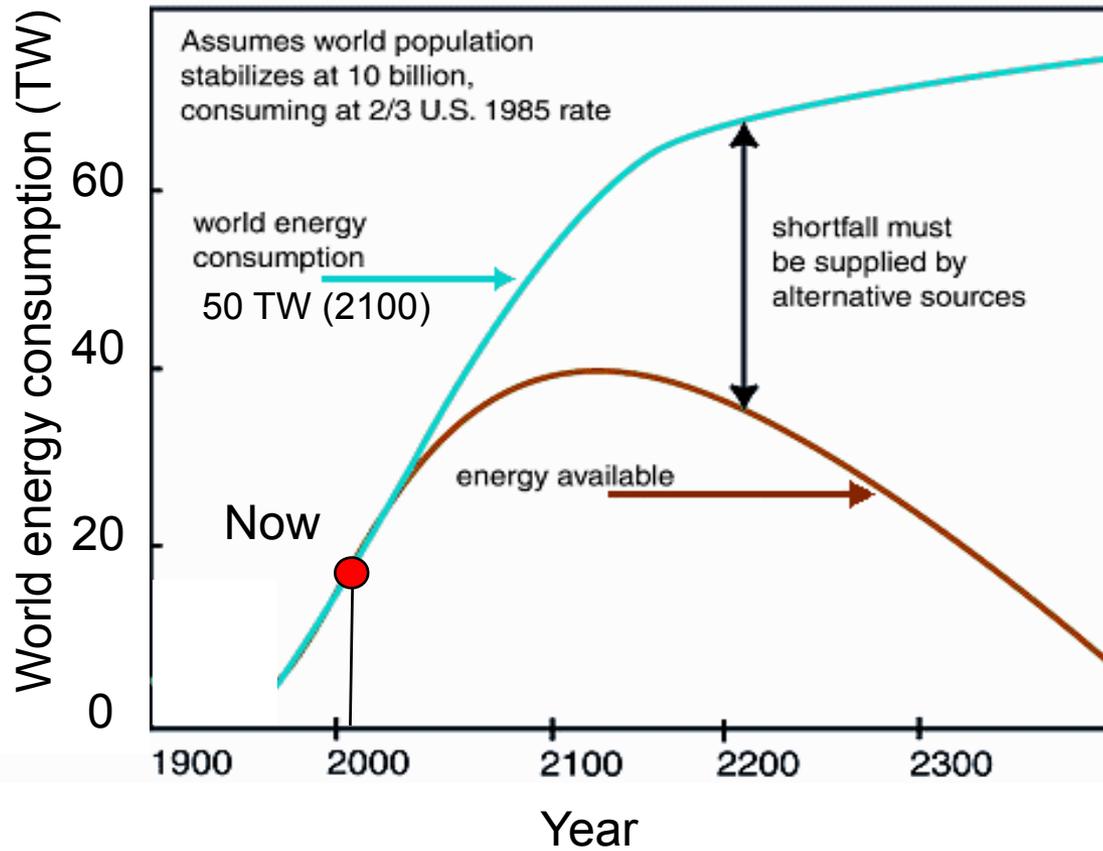
Possible Temperature Change (Relative to Pre-Industrial averages)



Europe
“20 20 by 2020”
Cut 20% of CO₂ by 2020, and produce 20% of electricity from renewables.

World energy outlook

Even if climate change was untrue (or if CO₂ sequestration worked fine) there is still a major issue regarding available sources (reserves of fossil fuels).



What are the options to fossil fuels?

Nuclear fission

- At present covers less than 7% of energy demand
- **Negligible CO₂ release** → renaissance
(notice fission saved more CO₂ than Kyoto protocol aims at)
- **Potential for a long-term solution**
(however not with the current once-through fuel cycle)
- at present the spent fuel is long-term highly active
- proliferation – danger of abuse of fissile materials
- public concerns regarding safety of the operation



NPP Temelin, PWR 2x1 GWe

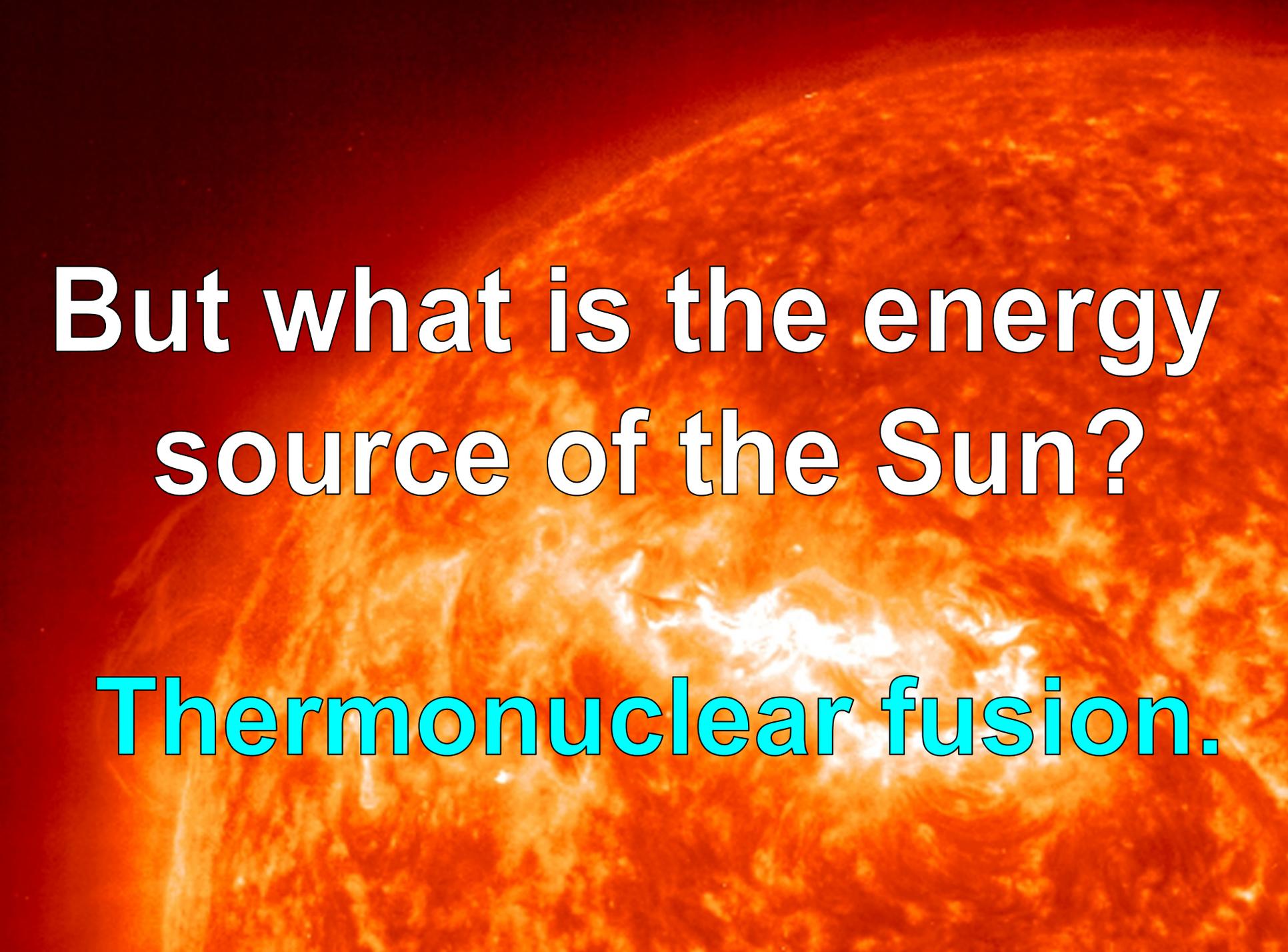


NPP Dukovany, PWR 4x450 MWe

Renewables

- ever increasing importance, at present renewables cover approx. 7% of the world energy demand
- low power density and poor reliability – the power grid still requires major backup (peaking) power plants and base load plants
- renewables not immune to the „NIMBY“ and geopolitical problems
- Sufficient potential available only in the solar power but our present knowledge is too limited to exploit it efficiently





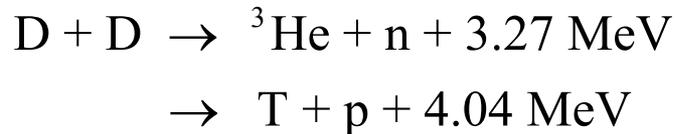
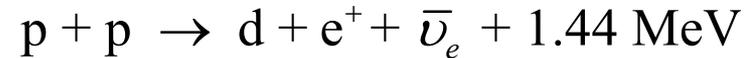
**But what is the energy
source of the Sun?**

Thermonuclear fusion.

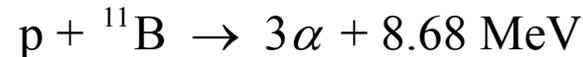
What is it, fusion?

By definition: **nuclear fusion is an exoergic reaction of light nuclei.**

Some very important examples:



In brief: $D(d,n){}^3\text{He}$ and $D(d,p)T$



What is it, thermonuclear fusion?

In fusion reactions, the two reacting nuclei are subject to long distance electric repulsion. This “**Coulomb barrier**” is overruled by much powerful “**strong interaction**” only at a very short distance. High kinetic energy of the reacting nuclei is the typical vehicle to get over the barrier.

We speak of **thermonuclear** fusion if this high kinetic energy is due to high temperature of the fuel.

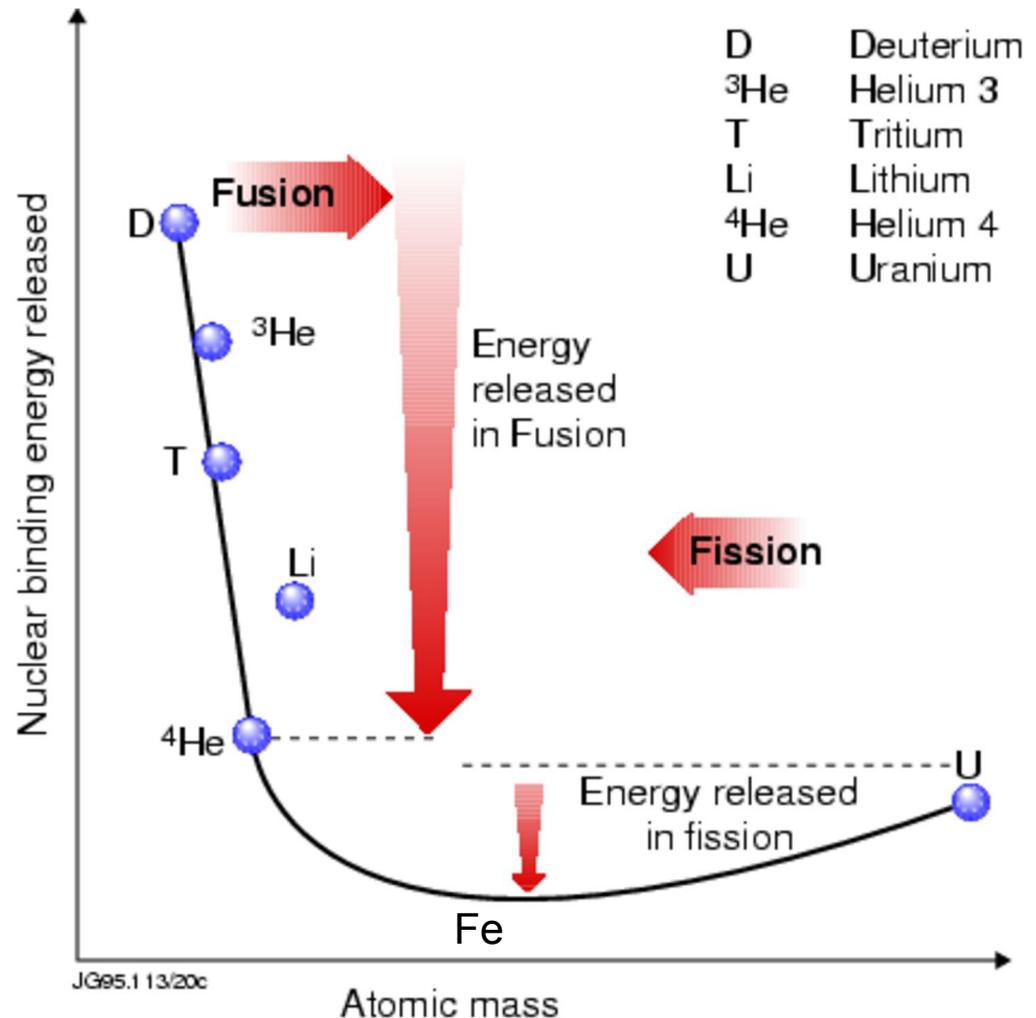
Both fusion and fission release energy

How comes?

Simple explanation is given by **liquid drop model** of nucleus

- strong nuclear interaction acts at a short range. It is therefore advantageous to form a sphere (to minimise surface)
- electrostatic interaction act on distance. Not only it causes a barrier to fusion, but also it means that for large nuclei it is advantageous to split (protons repulse each other)
- extremes like He are beyond this model, nuclear shell model required.

...but fusion is (even more) efficient



A more detailed look on the D-T fusion reaction

D-T fusion reaction has by far the highest probability

The diagram illustrates the D-T fusion reaction. On the left, a deuterium nucleus (D, one red proton and one blue neutron) and a tritium nucleus (T, one red proton and two blue neutrons) are shown colliding. An arrow points to the right, where the products are shown: a helium nucleus (He, two red protons and two blue neutrons) and a neutron (n, one blue neutron). The helium nucleus is labeled '20%' and the neutron is labeled '80%'. The reaction is summarized as $D + T \rightarrow {}^4\text{He} + n + \text{Energy}$.

Deuterium is a natural isotope of hydrogen (0.015 % of H)

Tritium has to be produced in the reactor from lithium

Helium (the alfa particle) carries one fifth of the released energy (2.5 MeV). It is charged i.e. confined \rightarrow it heats the plasma.

Neutron has no electric charge and carries four fifths of the released energy (14.1 MeV)

„Tritium breeding“ to be demonstrated and detailed in ITER

${}^6\text{Li} + n \rightarrow \text{T} + {}^4\text{He} + 7.7 \times 10^{-13} \text{ Joules}$

${}^7\text{Li} + n \rightarrow \text{T} + {}^4\text{He} + n - 4.0 \times 10^{-13} \text{ Joules}$

Famous quotes

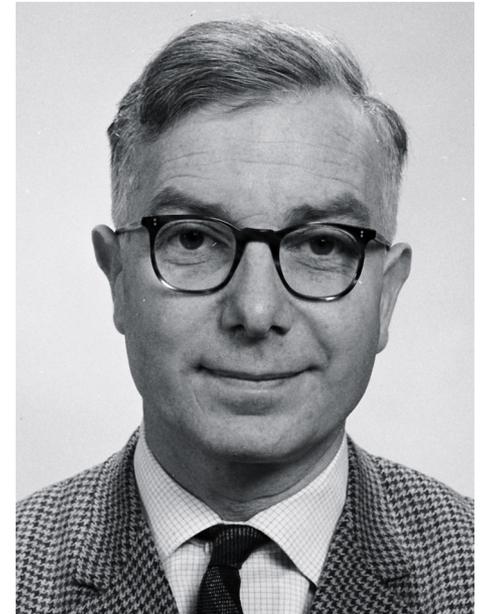


Sir Arthur S Eddington (1920):

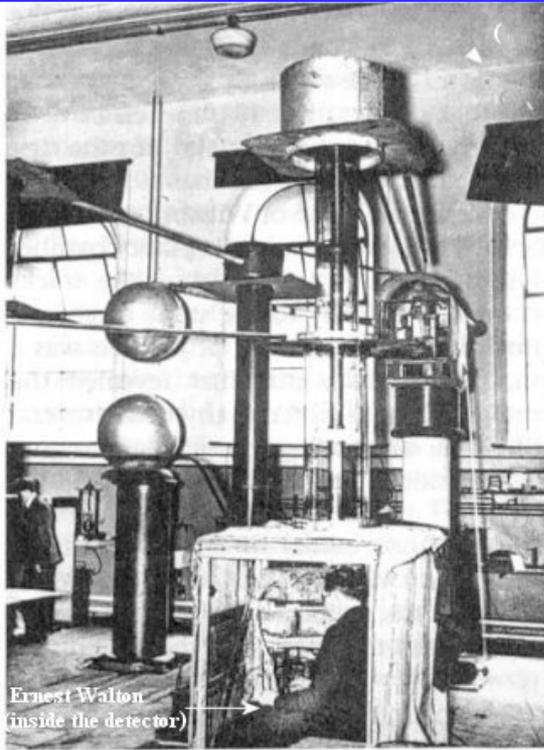
« If, indeed, the sub-atomic energy in the stars is being freely used to maintain their great furnaces, it seems to bring a little nearer to fulfilment our dream of controlling this latent power for the well-being of the human race – or for its suicide. »

John D. Lawson (1955):

« It is seen that for a useful reactor T must exceed 10^8 degrees and nt must exceed 10^{16} . These conditions are very severe. Conditions for a T-D-Li reactor (...) are easier though still severe. The corresponding values of temperature and nt are $T=3 \times 10^7$ degrees, $nt=10^{14}$. »



A brief historical note



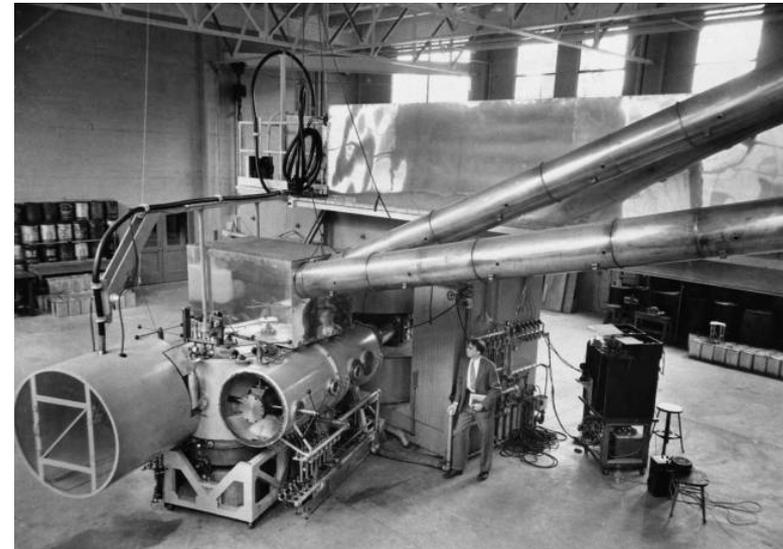
The first nuclear reaction 1932, Cockroft & Walton, 100 keV accelerator in the Cavendish laboratory
 ${}^7\text{Li}(p,\alpha)\alpha$ and ${}^6\text{Li}(p,\alpha){}^3\text{He}$ → Nobel prize 1951

D-D fusion discovered (!) 1934, Oliphant, Harteck, Rutherford, Cavendish laboratory...
 $\text{D}(d,p)\text{T}$ and $\text{D}(d,n){}^3\text{He}$

Both T and ${}^3\text{He}$ isotopes were isolated and tested by Alvarez in

Berkeley (first cyclotron) only in 1939
→ Surprise that ${}^3\text{He}$ is stable and T is not
→ Nobel prize 1968

Properties of the D-T fusion were studied for the first time during the World War II in the Purdue University. Large (resonant) cross-section came as a surprise.



Energy density in different fuels

Energy Source	Energy Density (MJ/kg)
Deuterium-Tritium Fusion (breeding Tritium out of Lithium)	187,000,000
Natural uranium (99.3% U-238, 0.7% U-235) in fast breeder reactor	24,000,000
Enriched uranium (3.5% U235) in light water reactor	3,456,000
Natural gas, Crude Oil and Derivates, Coal and similar (burned in air)	~40

Current fusion research aims at converting a breakthrough fraction of rest energy into high quality heat (that we can convert into electricity or fuel)

Conclusion of part I

In near future there is no magic solution to resolve increasing mismatch in energy demand and sustainable supply. Therefore

- (i) a broad mix of power sources must be maintained
- (ii) None of the potential sources can be ignored

